

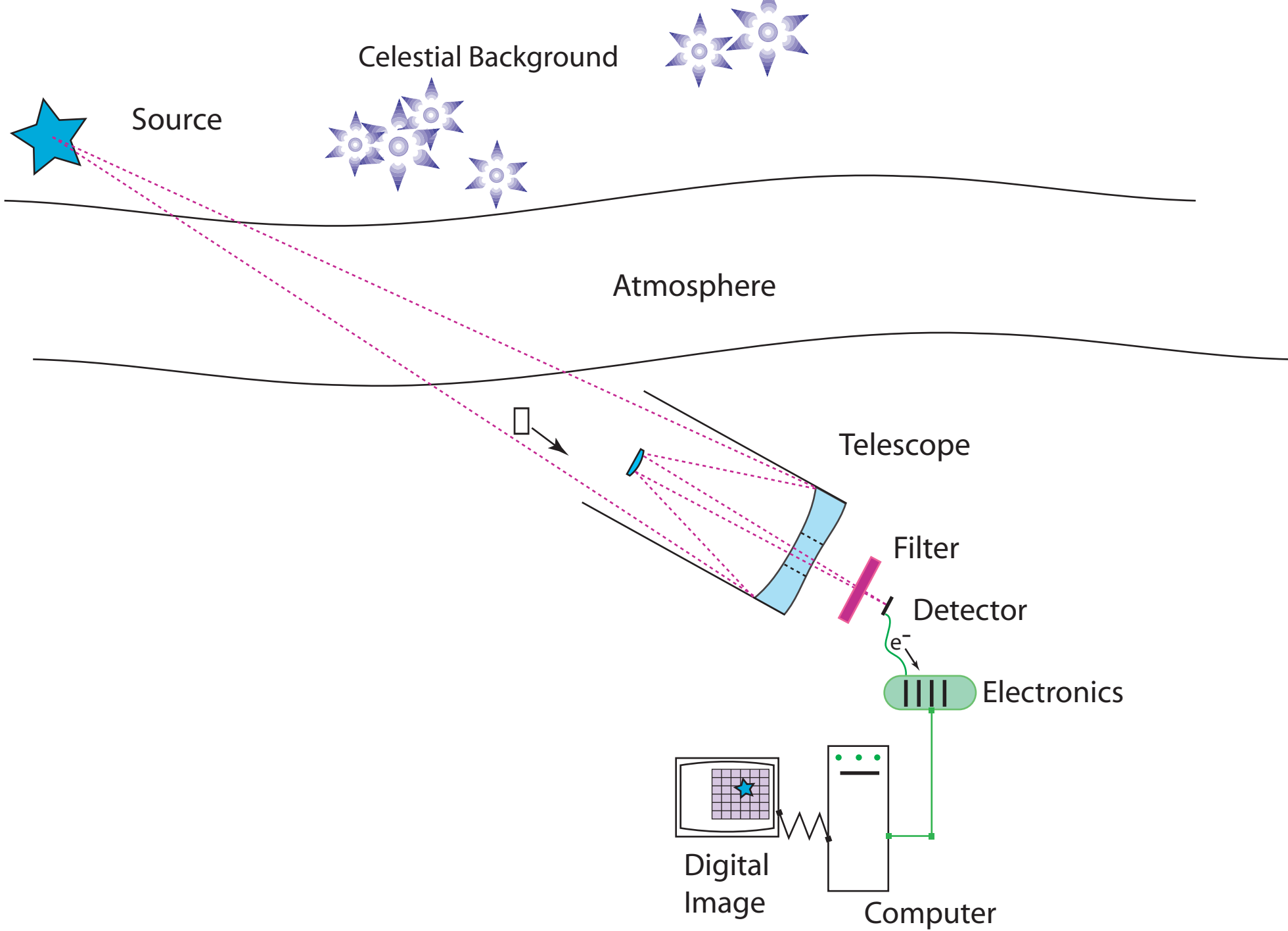
# Supernova Photometry - Lecture 1

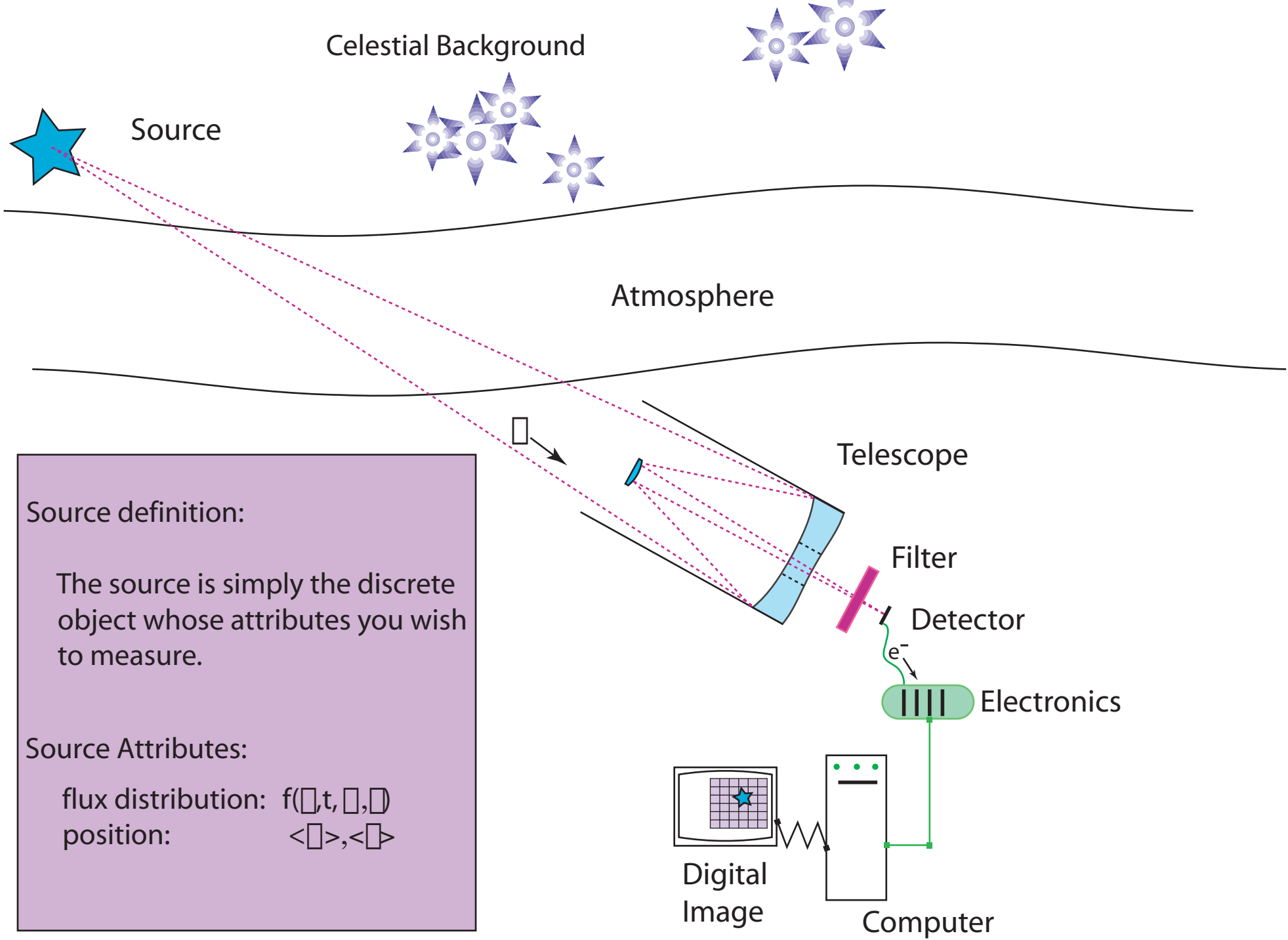
## The Photometry Signal Chain

LBL Photometry Group Meeting

Feb 9, 2004

presented by Greg Aldering

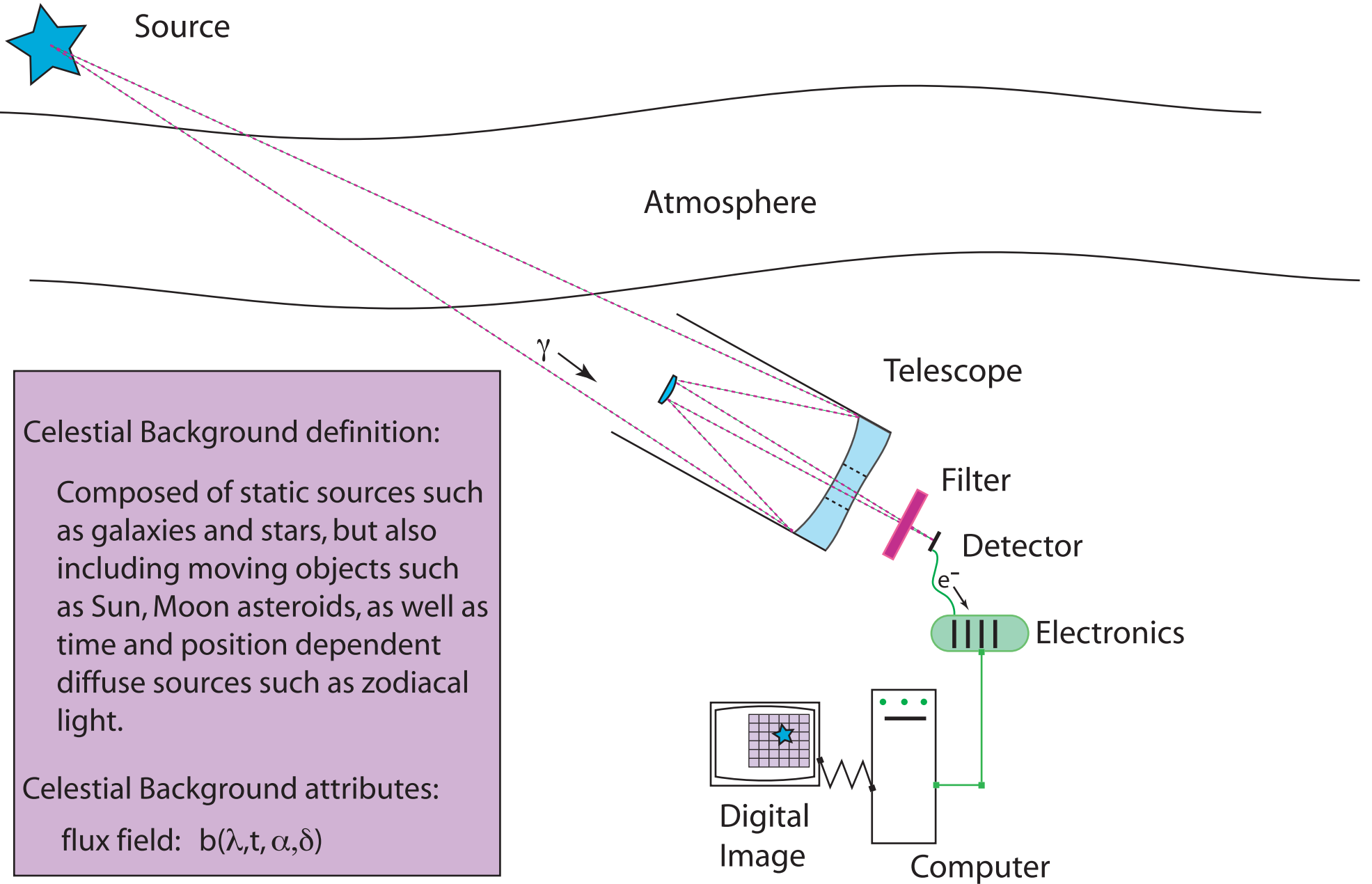




Source definition:  
 The source is simply the discrete object whose attributes you wish to measure.

Source Attributes:  
 flux distribution:  $f(x, t, y, z)$   
 position:  $\langle x \rangle, \langle y \rangle$

# Celestial Background

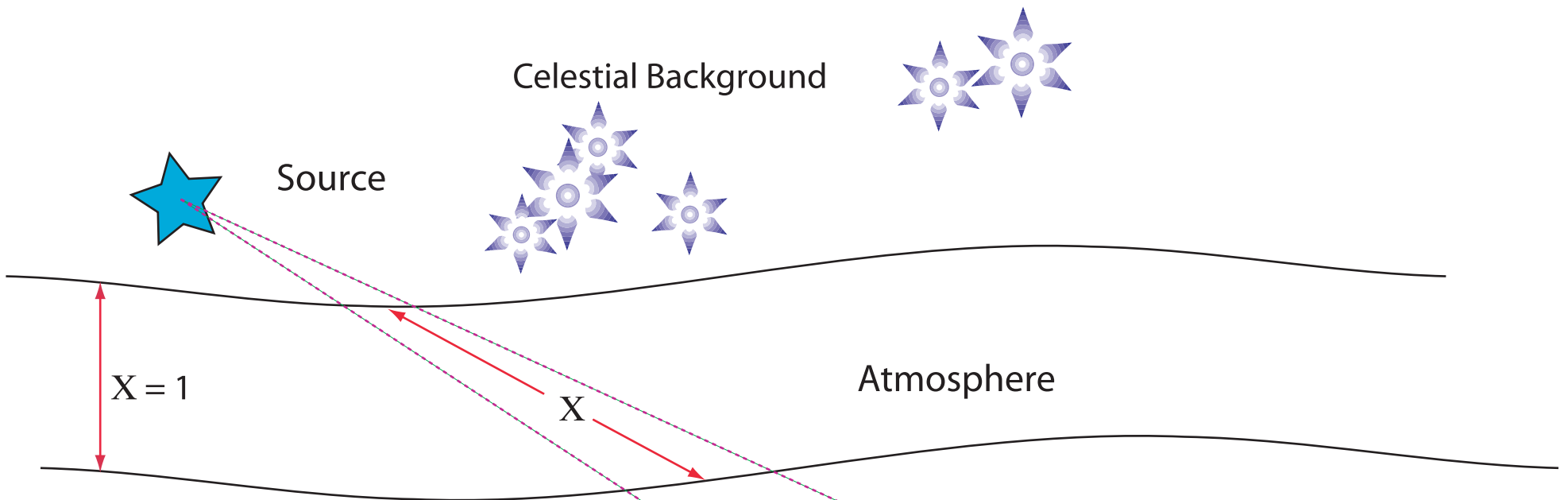


## Celestial Background definition:

Composed of static sources such as galaxies and stars, but also including moving objects such as Sun, Moon asteroids, as well as time and position dependent diffuse sources such as zodiacal light.

## Celestial Background attributes:

flux field:  $b(\lambda, t, \alpha, \delta)$



Atmosphere definition:

Atomic, molecular, crystalline, and particulate medium which absorbs, refracts and scatters light from the source & celestial background, as well as emitting light as a result of photo- and collisional-excitation and chemical processes.

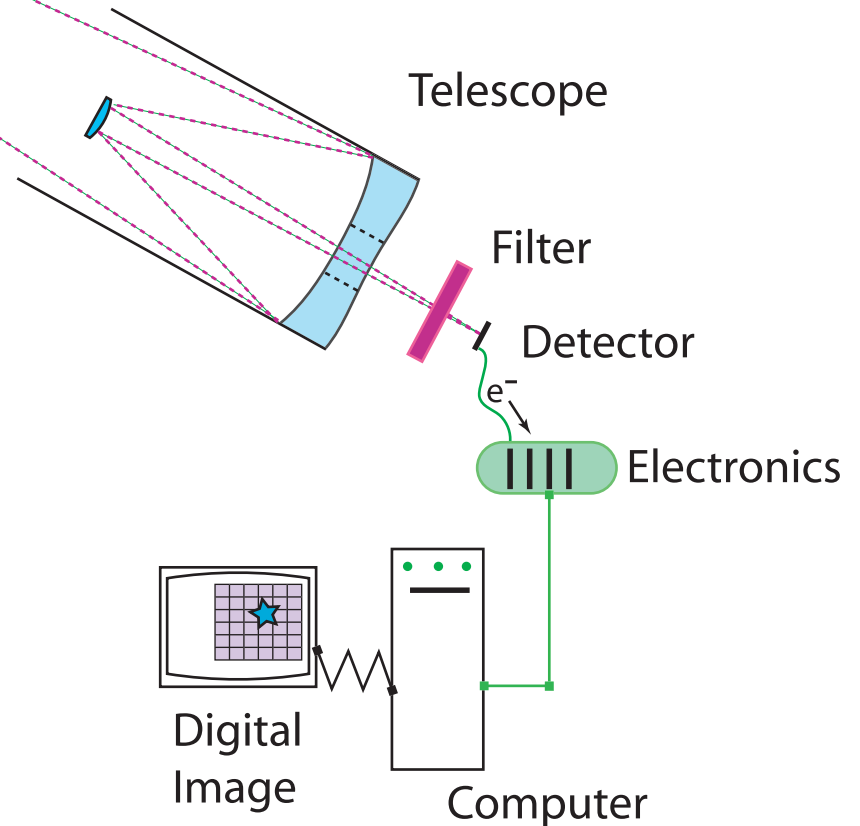
Atmosphere attributes:

extinction:  $k(\lambda, t, \alpha, \delta) \stackrel{?}{=} k(\lambda)$

airmass: X

brightness:  $s(\lambda, t, \alpha, \delta)$

distortion:  $MTF(\lambda, t, \omega) \leftrightarrow PSF(\lambda, t, \varpi)$



Celestial Background

Source

Atmosphere

Telescope

Filter

Detector

Electronics

Digital Image

Computer

Telescope definition:

Optic which collects light and magnify angles. Usually reflecting parabola, sometimes with refractive correcting optics.

Telescope attributes:

aperture:  $D$

focal length:  $f$

focal ratio:  $f/D$

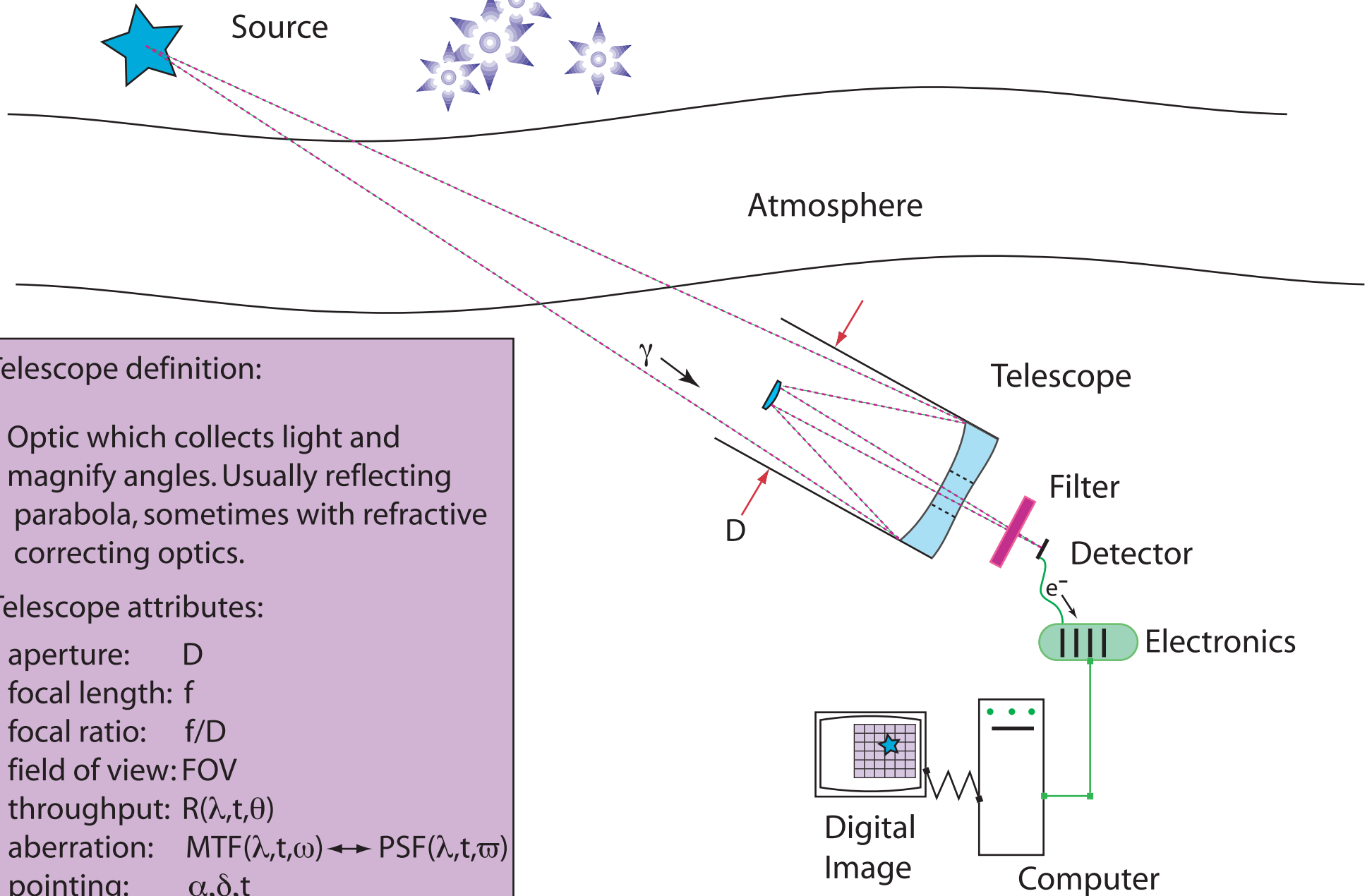
field of view: FOV

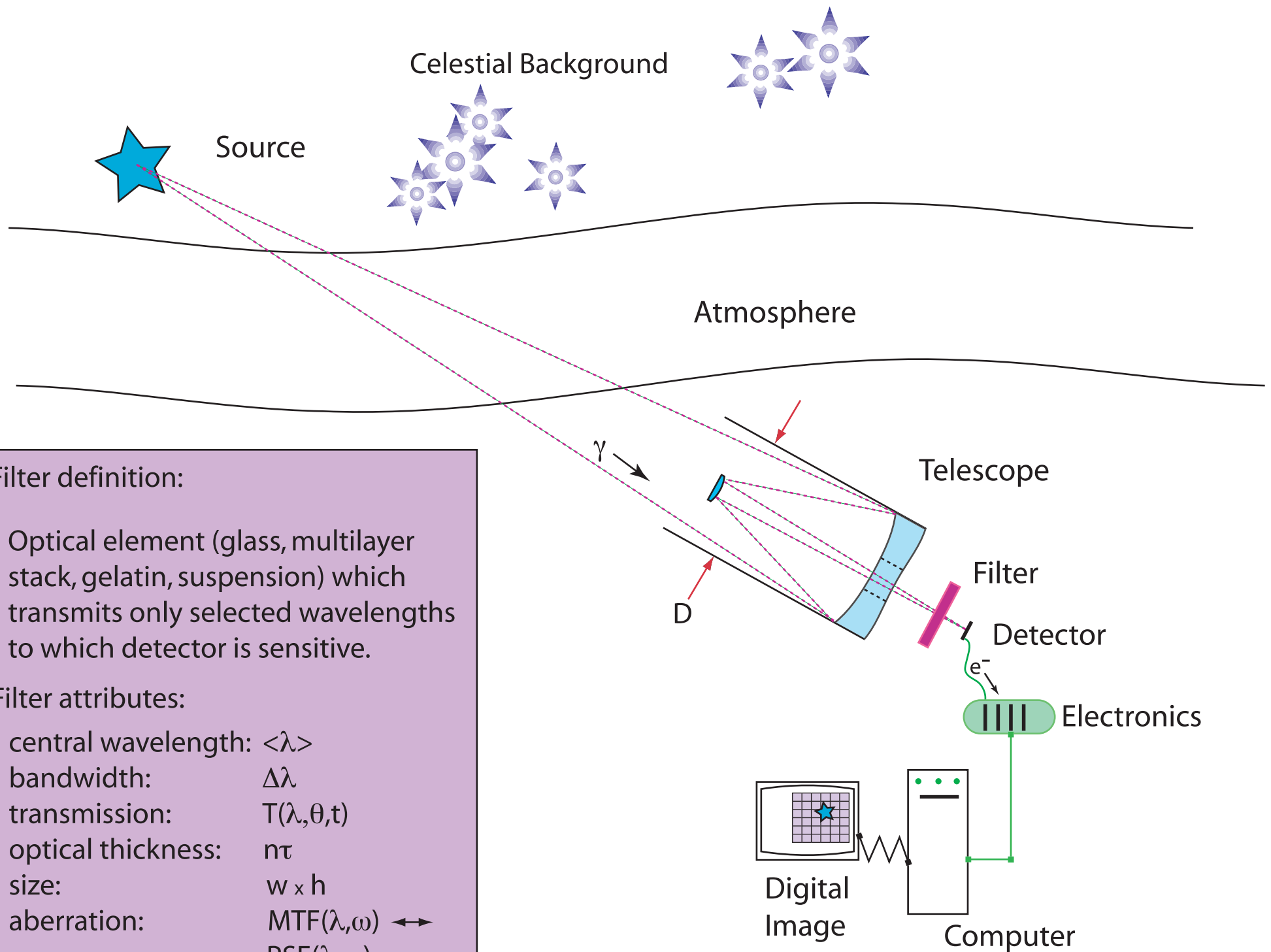
throughput:  $R(\lambda, t, \theta)$

aberration:  $MTF(\lambda, t, \omega) \leftrightarrow PSF(\lambda, t, \varpi)$

pointing:  $\alpha, \delta, t$

distortion:  $Jacobian(\alpha, \delta, x, y)$





Filter definition:

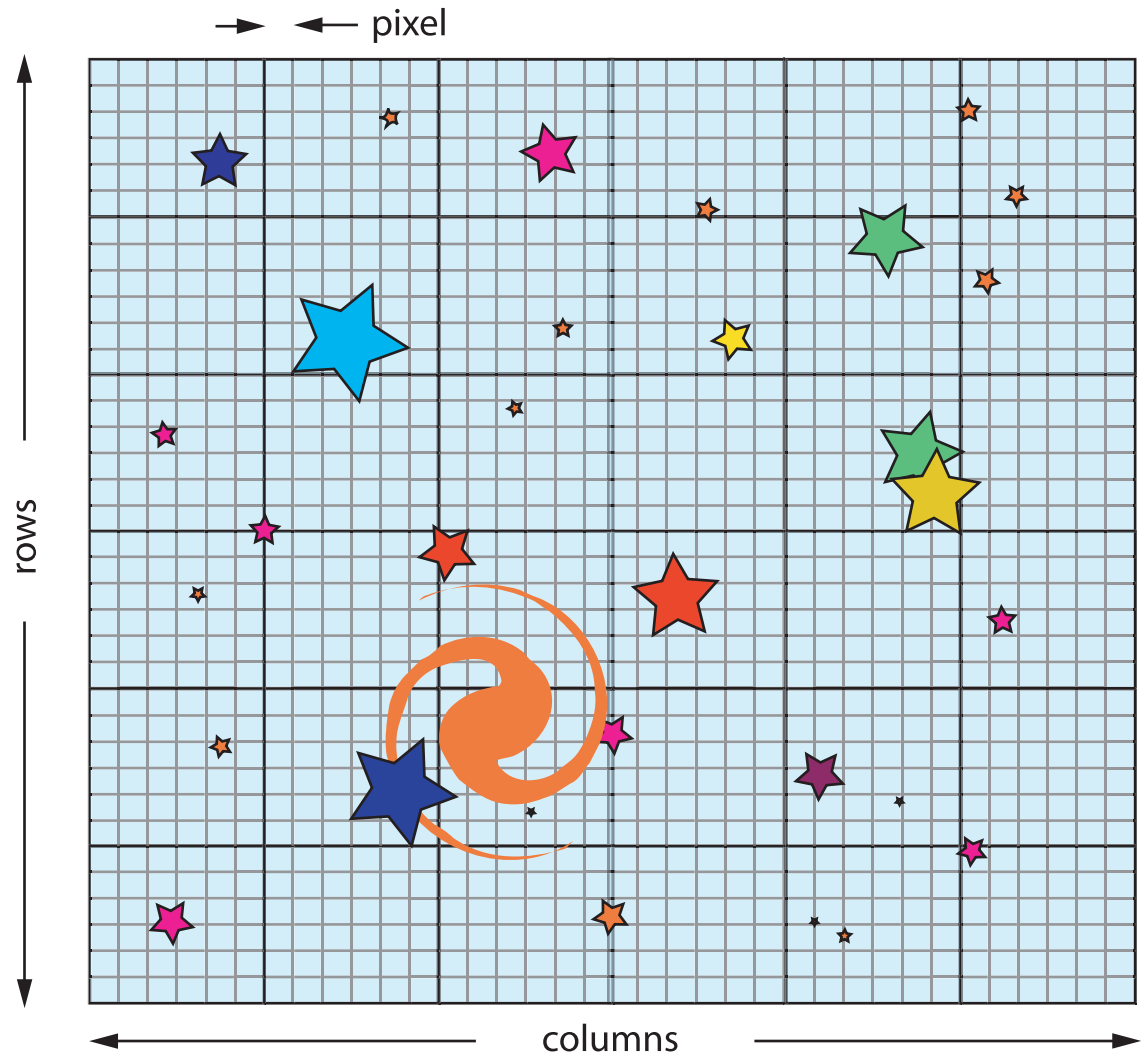
Optical element (glass, multilayer stack, gelatin, suspension) which transmits only selected wavelengths to which detector is sensitive.

Filter attributes:

central wavelength:  $\langle \lambda \rangle$   
 bandwidth:  $\Delta \lambda$   
 transmission:  $T(\lambda, \theta, t)$   
 optical thickness:  $n\tau$   
 size:  $w \times h$   
 aberration:  $MTF(\lambda, \omega) \leftrightarrow PSF(\lambda, \varpi)$

### CCD detector attributes:

efficiency:	$QE(\lambda, r, c)$
dark current:	$D(T, r, c)$
bias:	$B(r, c)$
fringing:	$F(\lambda, r, c)$
charge transfer:	CTE
gain:	$\mu\text{V}/e^-$
readout noise:	RON
well depth:	$C_{\text{max}}$
diffusion:	$MTF(x', y')$
persistence:	$C(t/\tau)$
amplifiers:	usually 1-4
defects:	bad columns, hot pixels, traps





## Electronics attributes:

gain: ADU/ $\mu$ V  
A/D range: e.g., 16 bits  
readout speed: e.g., 100 kpix/s  
noise/pickup: e.g., 60 Hz  
ghosts: response to bright  
object in another amp

## Computer attributes:

data format: e.g., FITS  
data scaling: e.g., BSCALE, BZERO  
data type: e.g. raw data are signed/unsigned 16 bit  
while processed data are 32 floating  
or truncated 16 bit integer