

Cosmic Microwave Background Radiation

After end of e^+e^- creation-annihilation leave the **lepton era**;
since neutrinos decouple about the same time (also nucleosynthesis),
enter **photon era**.

Photon gas in thermal equilibrium from $X\bar{X} \rightarrow \gamma\gamma$ and remains so
from **free-free** and **bremsstrahlung**: $e\gamma \rightarrow e\gamma$.

Processes freeze out at $z \approx 10^6$ ($t \approx 1500$ y), leaving only equilibrium
scattering (no energy change), so spectrum fixed by **thermalization epoch**
= **cosmic photosphere** (analogy to star).

Spectrum

Thermal equilibrium \Rightarrow **Planck** or **blackbody** spectrum

$$I_\nu = \frac{2h\nu^3}{c^2} \frac{1}{e^{h\nu/kT} - 1},$$

characterized by temperature T , with energy density ρ , number density n ,

$$\rho = \frac{\pi^2}{15} \frac{k^4}{c^5 \hbar^3} T^4 = 4.7 \times 10^{-34} (T/2.73 \text{ K})^4 \text{ g cm}^{-3}$$

$$n = \frac{2\zeta(3)}{\pi^2} \frac{k^3}{c^3 \hbar^3} T^3 = 410 (T/2.73 \text{ K})^3 \text{ cm}^{-3}$$

$$\nu_{peak} = 160 (T/2.73 \text{ K}) \text{ GHz}$$

$$\lambda_{peak} = 1.9 (T/2.73 \text{ K})^{-1} \text{ mm}$$

With expansion $\nu \sim a^{-1}$, $T \sim a^{-1}$ so spectrum remains blackbody.

Peaks today in microwave region.

Spectral distortion

Arise if out of equilibrium energy inputs (e.g. decaying particles, explosions). At $z > 4000$ energy shared by many photons so photon number \approx conserved \rightarrow like chemical potential:

$$\text{Bose - Einstein distortion : } \mu = 1.4\Delta E/E < 9 \times 10^{-5}.$$

At $z < 4000$ no time to Compton cool ($e\gamma \rightarrow e\gamma$) so get

$$\text{Comptonization : } \Delta E/E = e^{4y} - 1 \approx 4y$$

$$y = \int \Gamma dt = \sigma_T \int dt n_e(T_e/m_e) < 1.5 \times 10^{-5}$$

FIGURE of Planck and distortions

FIGURE of FIRAS spectrum

Last Scattering Surface

After decoupling ($\Gamma_{e\gamma} < H$) photons **free stream** since last interactions (electron scattering) done.

Not instantaneous though. **Optical depth** from scattering

$$\tau = \int_t^{t_0} n_e \sigma_T dt.$$

Since Poisson process then probability for scatter is $1 - e^{-\tau}$.

Visibility function gives probability of scattering between $z, z + dz$

$$p(z) = \frac{d}{dz}(1 - e^{-\tau}) = e^{-\tau} \frac{d\tau}{dz},$$

defining the last scattering surface. Well fit by Gaussian with mean $z_{lss} = 1055$ and thickness $\Delta z = 150$.

FIGURE 7.1 showing X, tau, p vs. z

Summary of Recombination Epochs

- Equality: energy density of matter and radiation equal; transition from radiation to matter dominated expansion behavior

- Recombination: (hydrogen) atoms form as ionization-recombination reactions fall out of equilibrium; free electron abundance freezes out
- Decoupling: electron-photon (Thomson) scattering reactions freeze out; matter and radiation lose contact
- Last Scattering: photons free stream

Recombination Epoch Parameters

$z_{eq} = 23900 \Omega_m h^2$	$t_{eq} = 1500(\Omega_m h^2)^{-3/2} \text{ y}$	$T_{eq} = 10 \text{ eV} = 10^5 \text{ K}$
$z_{rec} = 1300$	$t_{rec} = 2 \times 10^5 \text{ y}$	$T_{rec} = 0.3 \text{ eV} = 3600 \text{ K}$
$z_{dec} = 1100$	$t_{dec} = 3 \times 10^5 \text{ y}$	$T_{dec} = 0.25 \text{ eV} = 3000 \text{ K}$
$z_{lss} = 1055 \pm 75$		

Reionization

If electrons reheat after decoupling epoch ($z \approx 1100$) then get Compton distortion – probe of galaxy/star formation energetics.

$$y = 2.3 \times 10^{-5} \left(\frac{T_e}{5000 \text{ K}} \right) \left(\frac{z_{heat}}{1000} \right)^{3/2}.$$

Optical depth from scattering

$$\begin{aligned}\tau(z) &= \sigma_T \int_{t(z)}^{t_0} dt n_e = \sigma_T \int_0^z dz (dt/dz) n_\gamma (n_e/n_\gamma) \\ &= \sigma_T H_0^{-1} \eta n_{\gamma,0} \int dz (1+z)(1+\Omega z)^{-1/2} \\ &\rightarrow \left(\frac{1+z}{121}\right)^{3/2} h^{-1} \quad (\Omega = 1)\end{aligned}$$

So universe unlikely to have significant optical depth after decoupling.

(Local version called *Sunyaev-Zel'dovich effect*).

Cosmological Origin of Blackbody Radiation

- 5 times the energy density of galactic light
- Would require $10^9 n_{gal}$ in discrete sources to match isotropy
- Direct detection at $z > 0.54$ from SZ effect; indirect out to $z = 1.97$ by

atomic excitation

- Predicted by Big Bang; consistent with nucleosynthesis

Temperature, Expansion, and Velocity

Expansion gives scaling

$$T(z) = T_0 (1+z)$$

Observed to 13% at $z < 1.97$ (upper limits at $z < 4.38$) from radiative excitations of CI, CII, CO.

Velocity gives Doppler shift

$$T = T_0 (1 + v \cos \theta) \quad \text{dipole anisotropy}$$

Motion with respect to isotropic CMB frame. Detected

$$\Delta T/T = 1.23 \times 10^{-3} \quad \Delta T = 3.356 \text{ mK}$$

$$v_{\oplus} = 369 \text{ km s}^{-1} \quad v_{gal} = 550 \text{ km s}^{-1}$$

This motion gives probe of mass distribution. **Dipole offset** of CMB from local peculiar velocity field depends on $\Omega_m^{0.6}$.